

# Is the dust-to-gas ratio constant in molecular clouds?

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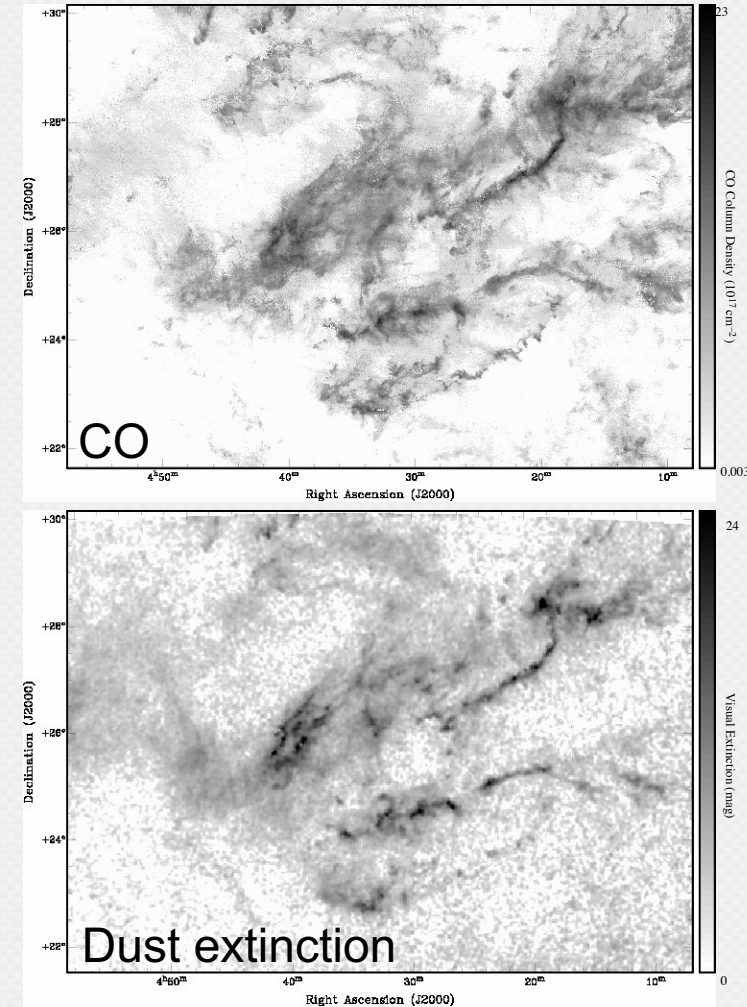






# Dust in Molecular Clouds

- Molecular hydrogen emits weakly, so bulk of gas is effectively invisible
- 1% dust-to-gas mass ratio used to infer gas mass in dense, cold molecular clouds
- Turn to tracers: other molecules (CO,  $\text{N}_2\text{H}^+$ ), and/or dust
- Dust seems to correlate well with gas



Pineda et al (2010)

# Dust in Molecular Clouds

But is the dust-to-gas ratio in molecular clouds uniformly 1%?

## •Anomalous extinction laws:

- $R_v = A_v / (A_b - A_v) = 3.1$  in diffuse clouds (fairly “universal”), but  $R_v \sim 5$  in molecular clouds (Cardelli+ 1989; Weingartner & Draine 2001)
- Implies molecular clouds have a different grain size distribution

## •Coreshine phenomenon:

- higher abundance of large grains ( $>1$  micron) in dense filaments (Pagani+ 2010; Steinacker+ 2010; Lefevre+ 2014)

## •Ophiuchus A:

- Spatial mismatch between gas and dust (Di Francesco+ 2004; Friesen+ 2014)
- mean 1.1% dust-to-gas ratio, but with local fluctuations of 0.5% to 10% (Liseau+ 2015)



# Dust is not Gas!

- Dust behaves dynamically different than gas
- Could supersonic turbulence cause dynamical variations in the dust-to-gas ratio?
- Analytic estimates for 0.1 micron dust grains in molecular clouds
- For a molecular cloud of typical values:
  - density  $\sim 10^{-20} \text{ g/cm}^3$ , sound speed  $\sim 0.2 \text{ km/s}$ ,
  - 0.1 micron grains, intrinsic grain density  $\sim 3 \text{ g/cm}^3$ ,
  - Drag stopping timescale  $\sim 10^3 \text{ yrs} \ll 10^6$  dynamical time of molecular cloud
  - Thus, expect a well-coupled mixture of dust and gas. High drag regime.
- But this assumes cloud is homogenous and *ignores turbulence!*

# Supersonic Dusty Turbulence



We modelled the turbulent dynamics of dusty molecular clouds using the SPH code Phantom ([Price+ 2017](#))

- Not concerned with grain growth or destruction
- No self-gravity (not trying to make stars) or magnetic fields

## Initial Conditions:

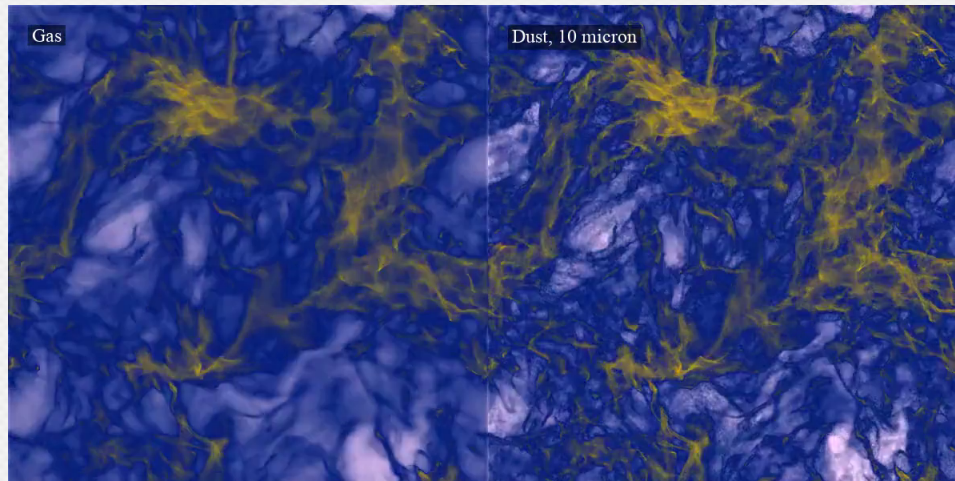
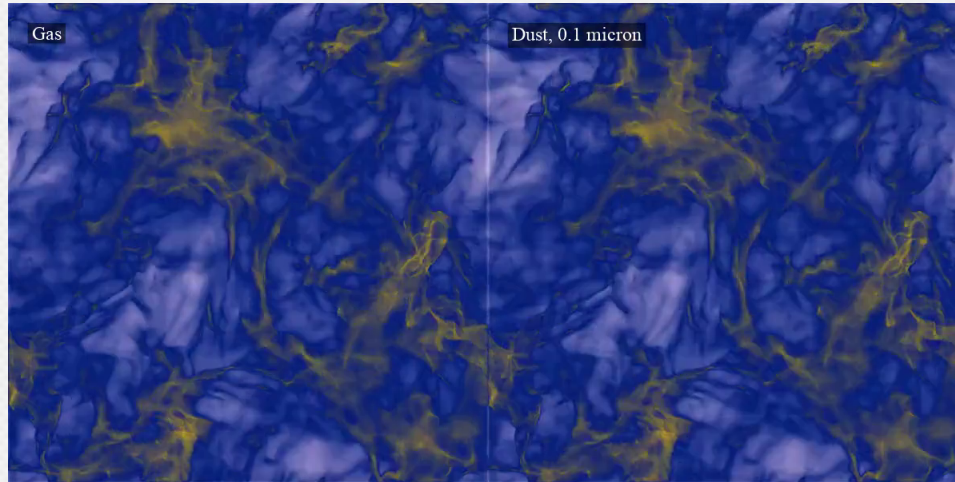
- Density =  $10^{-20}$  g/cm<sup>3</sup>
- Isothermal gas with  $T = 11.5$  K,  $c_s = 0.2$  km/s
- 3 parsec length box
- Driven turbulence on large scales, for  $\sim 14$  Myr

## Dust Physics:

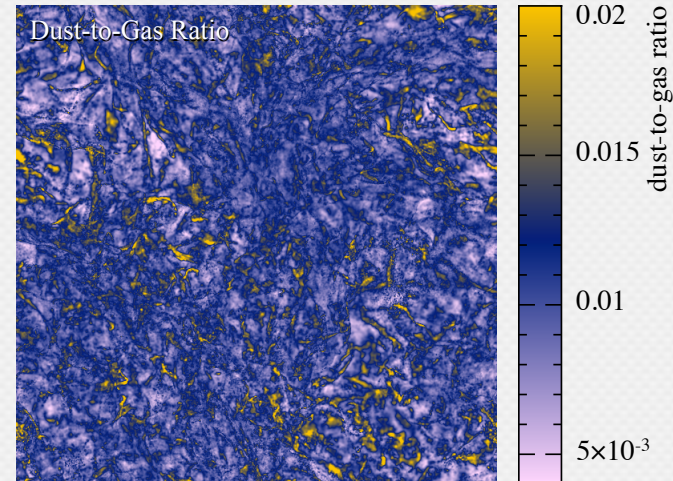
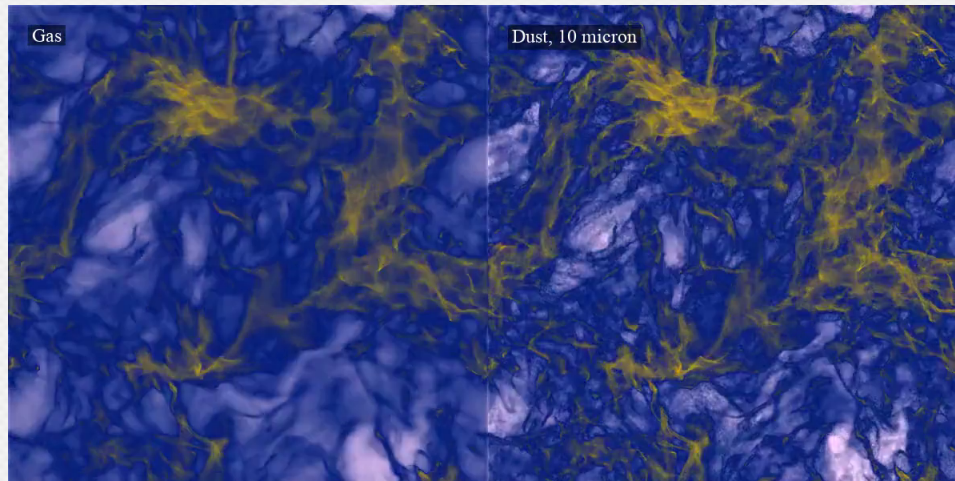
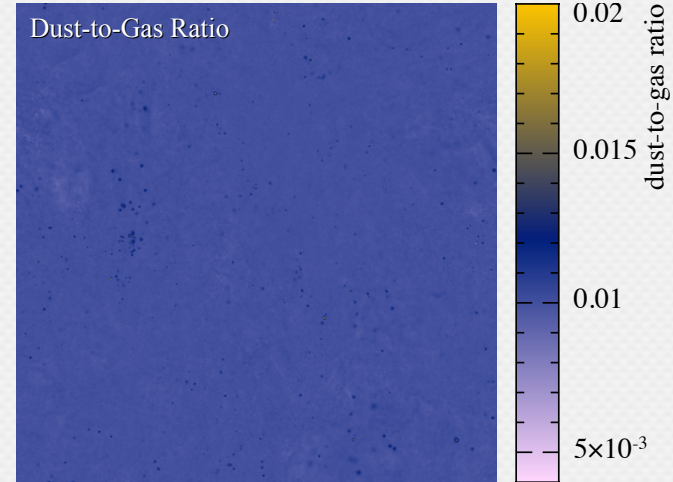
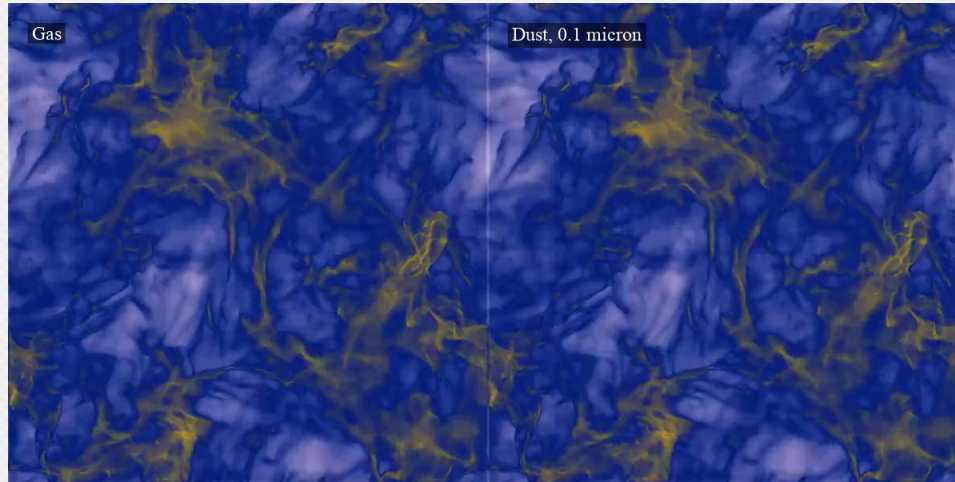
- 3 separate simulations for 0.1, 1 and 10 micron grain sizes
- Initially uniform 1% dust-to-gas ratio



# Column Dust and Gas Density



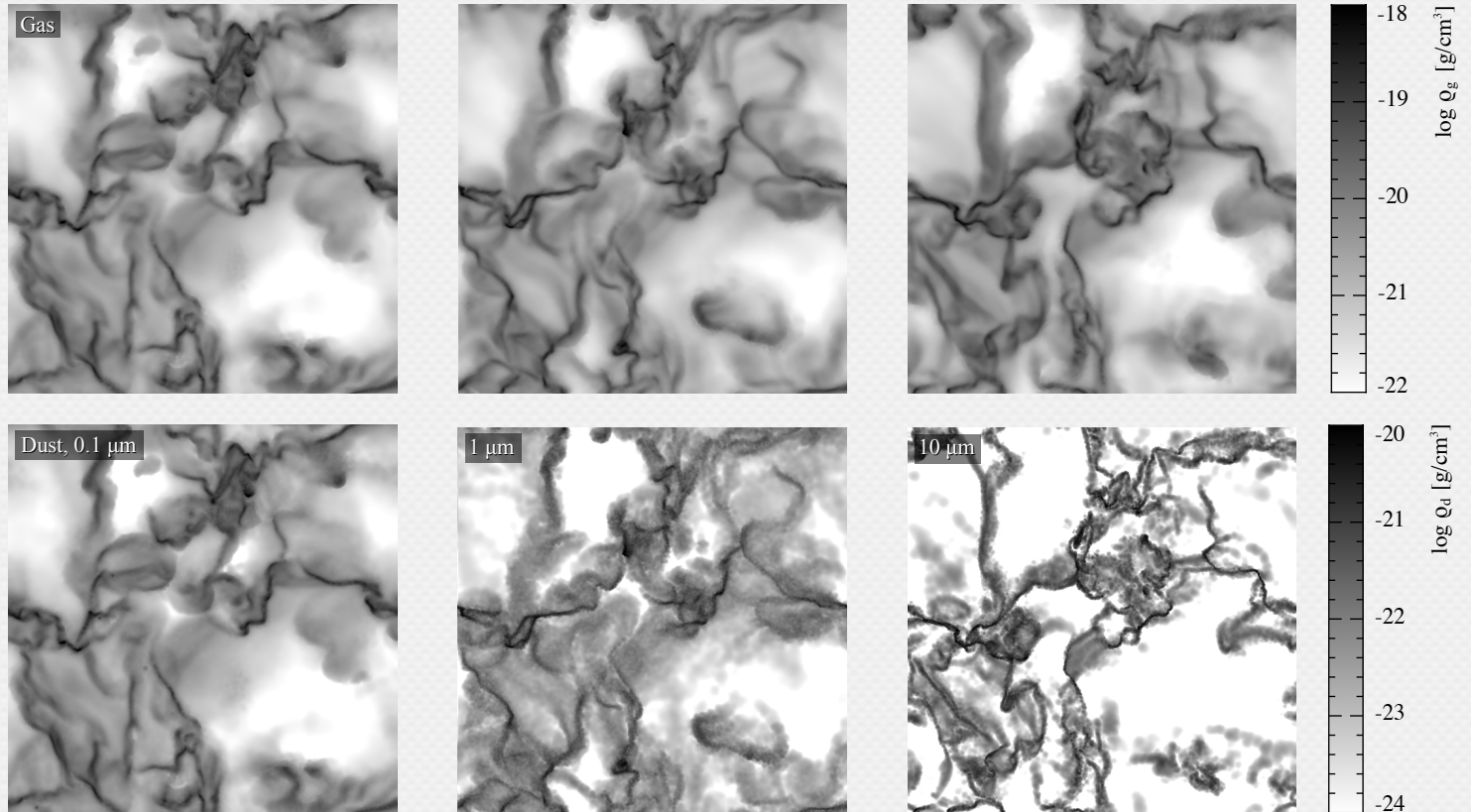
# Column Dust and Gas Density





# Size-sorting of Dust

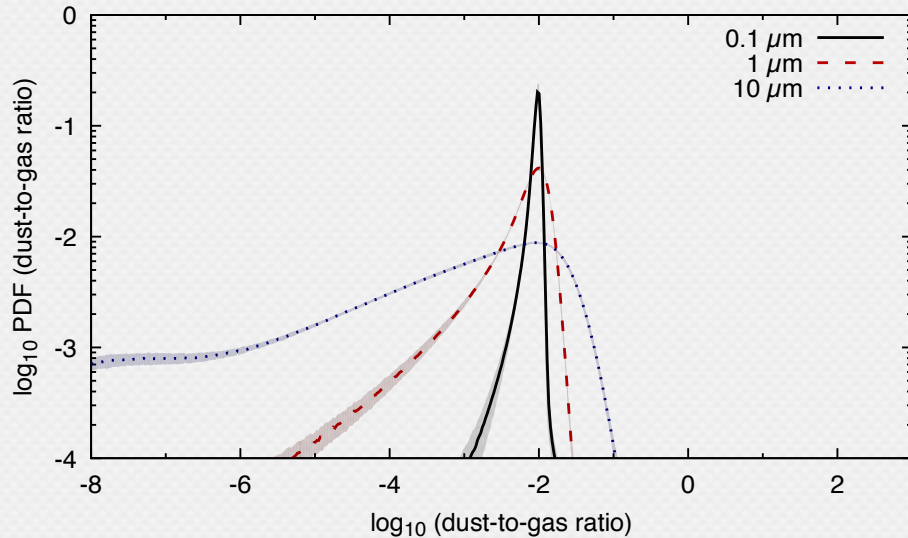
Slices of gas and dust through midplane of the cloud



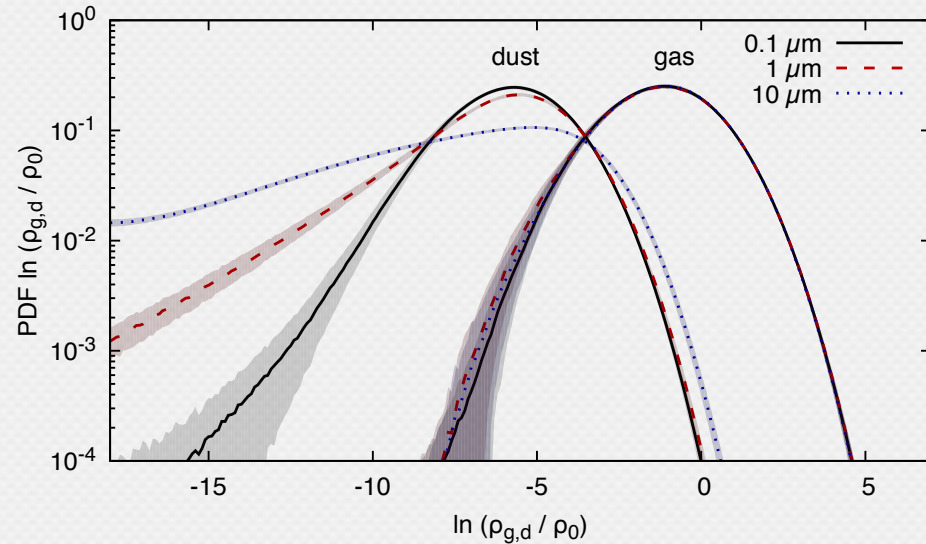
- 0.1 micron grains remain well-coupled to the gas throughout the cloud
- Large grains preferentially concentrated in filaments
- Up to 10x increase of 10 micron grains in dense filaments

# Probability Distributions

PDFs of dust-to-gas ratios



PDFs of gas and dust densities

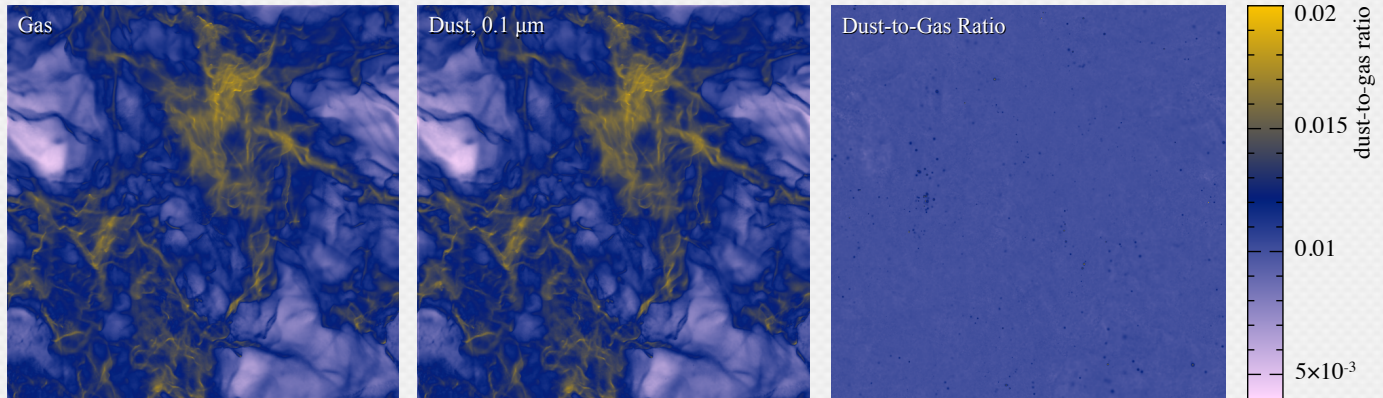


- 0.1 micron grains:
  - Sharply peaked PDF of dust-to-gas ratios at 1%
  - Matching log-normal distribution of gas and dust density
- 1-10 micron grains:
  - PDFs broaden due to 'size-sorting'
  - Turbulence causes dynamical transfer of dust mass into high density filaments

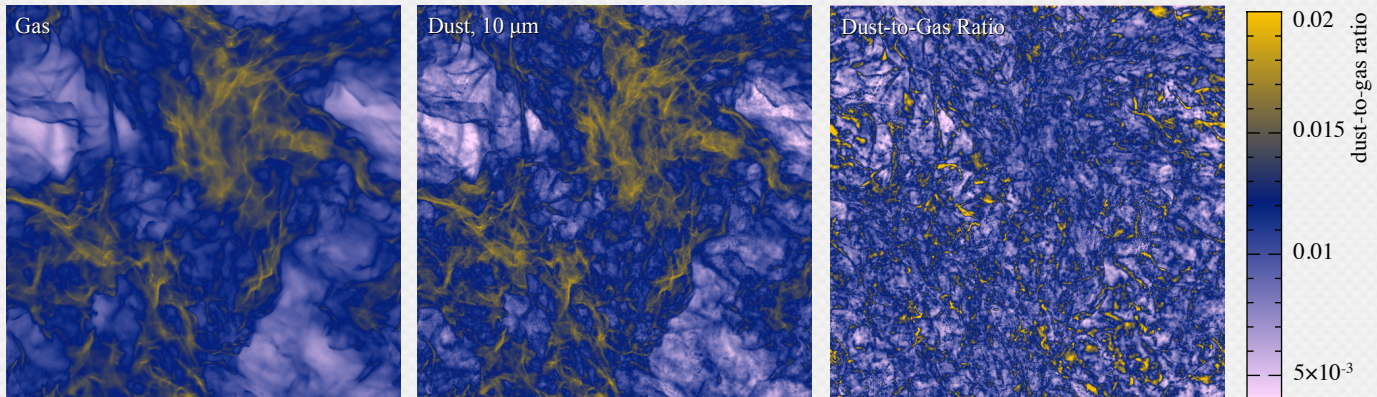


# Is the dust-to-gas ratio constant?

- **Yes!** For 0.1 micron grains, turbulence has almost no effect



- **No!** For  $\geq 1$  micron grains, turbulence can induce large variations



# Conclusions

Studied the dynamical effect of supersonic turbulence on distribution of dust in molecular clouds:

- 0.1 micron dust grains:
  - Minimal variation in the dust-to-gas ratio
  - Dust well-coupled to the gas throughout molecular cloud
- 1-10 micron grains:
  - ‘size-sorting’, concentration of large grains into filaments
  - Up to 10x increase of 10 micron grains in dense filaments
  - Related to coreshine?



# Dust Algorithm

## Low Drag: Two fluid approach

- Separate populations of dust and gas
- Evolve gas velocity, dust velocity, gas density, dust density
- Works well when they behave as separate fluids
  - have to resolve drag timescale and lengthscale

## High Drag (molecular clouds): One fluid approach

- Combined mixture of dust and gas
- Evolve barycentric velocity, combined density, dust-to-gas ratio
- Works well when they behave similarly
  - does not work well when they start behaving as separate fluids